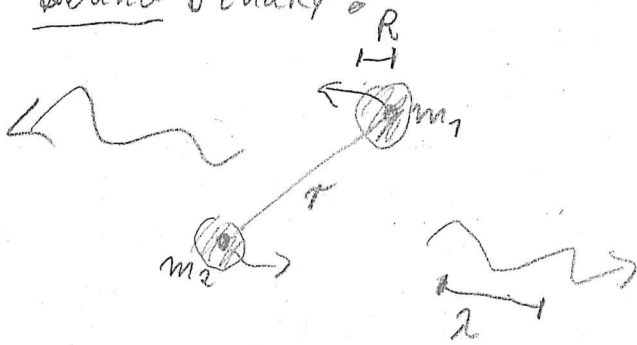


# post-Newtonian (PN) approximation

bound binary:



if  $R \ll r \ll \lambda$  (hierarchy of scales)  
then PN approx. possible.

expansion parameter:

$$\epsilon = \frac{v^2}{c^2} \sim \frac{GM}{c^2 r}$$

$\lambda \sim \frac{c}{\nu}$  3rd Kepler (bound orbits)

one PN order  $\approx \frac{1}{c^2}$

compact object  $\rightarrow$  multipole expansion in  $\frac{R}{r}$

proper time  
 $X_a(t_a), a=1,2$   
point particle (pp)

action:

$$S_a^{pp} = - \int dt_a m_a c^2 + \mathcal{O}\left(\frac{R^2}{r^2}\right)$$

point mass  $\rightarrow$  spin  $\rightarrow$  exercise (2g-coupling)  
MORE later...

PN approx. for orbital motion:

the Fokker action

starting point:  $S = S_E + S_1^{pp} + S_2^{pp}$

Einstein action

field eq. for  $g_{\mu\nu}$  from  $S$ :  $\delta S = \frac{\delta S}{\delta g_{\mu\nu}} \delta g_{\mu\nu}$

solve perturbatively in  $\epsilon \sim \frac{v^2}{c^2} \stackrel{!}{=} 0$  (\*)

$\rightarrow g_{\mu\nu}[X_1, X_2]$

insert solution into  $S \rightarrow S'[X_a^\mu] = S[g_{\mu\nu}[X_a], X_a^\mu]$  (Fokker action)

then:  $\delta S' = \left( \frac{\delta S}{\delta g_{\mu\nu}} \Big|_{g=g[X_a]} \cdot \frac{\delta g_{\mu\nu}[X_a]}{\delta X_a^\mu} + \frac{\delta S}{\delta X_a^\mu} \right) \delta X_a^\mu$

0 since  $g_{\mu\nu}[X_a]$  satisfies (\*)

$\sim \frac{\delta S'}{\delta X_a^\mu} = \frac{\delta S}{\delta X_a^\mu} \stackrel{!}{=} 0 \rightarrow$  EOM for  $X_a$  unchanged!

[Remark: mistakes in literature]

Let's calculate to 1PN:

ansatz for  $g_{\mu\nu}$ :

$$c^2 dx^2 = -g_{\mu\nu} dx^\mu dx^\nu$$

$$= e^{2\phi/c^2} \left( c dt - \frac{1}{c^2} A_i dx^i \right)^2 - e^{-2\phi/c^2} \left( \delta_{ij} + \frac{1}{c^4} G_{ij} \right) dx^i dx^j$$

"PN fields"  $\phi, A_i, G_{ij}$   
 ↑ ↑  
 gravito-electric gravito-magnetic

point mass:

$$S_1^{pp} = -m_1 c^2 \int dt \frac{dx}{dt} + \mathcal{O}\left(\frac{R^2}{r^2}\right)$$

$$V_1^i = \frac{dx_1^i}{dt}$$

$$= \int dt m_1 \left[ -c^2 + \frac{1}{2} V_1^2 - \phi + \frac{1}{c^2} \left( \frac{1}{8} V_1^4 - \frac{3}{2} V_1^2 \phi - \frac{1}{2} \phi^2 + A_i V_1^i \right) + \mathcal{O}(c^{-4}) \right]$$

(exercise: include spin  $\rightarrow$  gravito-magnetic dipole)

$\vec{x} = \vec{x}_a$

Einstein action:

$$S_E = \frac{c^4}{16\pi G} \int dt d^3x \sqrt{-g} R_{ij} \partial_i \partial_j$$

$$= \frac{1}{32\pi G} \int dt d^3x \left( 4\phi \Delta \phi - \frac{4}{c^2} \phi \partial_t^2 \phi - \frac{1}{c^2} A_i \Delta A_i + \mathcal{O}(c^{-4}) \right) + \text{boundary terms}$$

in Lorenz gauge  $\partial_\nu \bar{h}^{\mu\nu} = 0 \Leftrightarrow \partial_i A_i + 4\partial_t \phi = 0$   
 $\phi_{ij}$  is higher order

sort terms:

$$S = \int dt L_{kin} + \int dt d^3x \left( \gamma_\phi \phi - \frac{1}{2} \phi \mathcal{D}_\phi \phi + \frac{1}{c^2} \gamma_A^i A_i - \frac{1}{2c^2} A_i \mathcal{D}_A A_i \right) + \mathcal{O}(c^{-4})$$

$$L_{kin} = m_1 \left( -c^2 + \frac{V_1^2}{2} + \frac{V_1^4}{8c^2} + \dots \right) + (1 \leftrightarrow 2)$$

$$\gamma_\phi = -m_1 \left( 1 + \frac{3V_1^2}{2c^2} \right) \delta_1 + (1 \leftrightarrow 2)$$

where  $\delta_1 \equiv \delta(x - x_1(t))$

$\hookrightarrow 1 = \int d^3x \delta_1$

$$\mathcal{D}_\phi = -\frac{1}{4\pi G} (\Delta - c^{-2} \partial_t^2) + \frac{m_1}{c^2} \delta_1 + \frac{m_2}{c^2} \delta_2$$

$$\gamma_A^i = m_1 V_1^i \delta_1 + (1 \leftrightarrow 2), \quad \mathcal{D}_A = \frac{1}{16\pi G} \Delta$$

field eqs.: (vary  $\phi$  &  $A_i$ )

$$\mathcal{D}_\phi \phi = \gamma_\phi, \quad \mathcal{D}_A A^i = \gamma_A^i \quad (\text{exercise})$$

Solutions:

$$\phi = \mathcal{D}_\phi^{-1} \gamma_\phi, \quad A^i = \mathcal{D}_A^{-1} \gamma_A^i$$

insert into  $S \rightarrow S'$ :

$$S' = \int dt L_{kin} + \int dt d^3x \left( \underbrace{\gamma_\phi \mathcal{D}_\phi^{-1} \gamma_\phi}_{\frac{1}{2} \gamma_\phi \mathcal{D}_\phi^{-1} \gamma_\phi} - \frac{1}{2} (\mathcal{D}_\phi^{-1} \gamma_\phi) \mathcal{D}_\phi \mathcal{D}_\phi^{-1} \gamma_\phi + \frac{1}{2c^2} \gamma_A^i \mathcal{D}_A^{-1} \gamma_A^i \right) + \mathcal{O}(c^{-4})$$

evaluate A-term: (the last one)

with formula:  $\Delta^{-1} \delta_a = -\frac{1}{4\pi |\vec{x} - \vec{x}_a|}$

$$L_A = \int d^3x \frac{1}{2c^2} \gamma_A^i \mathcal{D}_A^{-1} \gamma_A^i = \int \frac{d^3x}{2c^2} (m_1 v_1^i \delta_1 + m_2 v_2^i) \left( \frac{\Delta}{16\pi G} \right)^{-1} (m_1 v_1^i \delta_1 + m_2 v_2^i \delta_2)$$

$$= \int d^3x \frac{8\pi G}{c^2} \left( \underbrace{m_1 v_1^i \delta_1 \Delta^{-1} m_1 v_1^i \delta_1}_{\sim \frac{1}{|\vec{x}_1 - \vec{x}_1|} \text{ divergent self-energy } \rightarrow \text{drop}} + \underbrace{m_1 v_1^i \delta_1 \Delta^{-1} m_2 v_2^i \delta_2 + m_2 v_2^i \delta_2 \Delta^{-1} m_1 v_1^i \delta_1}_{\text{part. int.}} + \underbrace{m_2 v_2^i \delta_2 \Delta^{-1} m_2 v_2^i \delta_2}_{\sim \frac{1}{|\vec{x}_2 - \vec{x}_2|} \rightarrow \text{drop}} \right)$$

where  $r = |\vec{x}_1 - \vec{x}_2|$

$$= -\frac{4G}{c^2 r} m_1 m_2 v_1^i v_2^i$$

$\phi$ -term:

$$L_\phi = \int d^3x \frac{1}{2} \gamma_\phi \mathcal{D}_\phi^{-1} \gamma_\phi$$

$2 \times 4 \times 2 = 16$  terms

$\rightarrow$  use diagrams to keep track

first: operator inverse

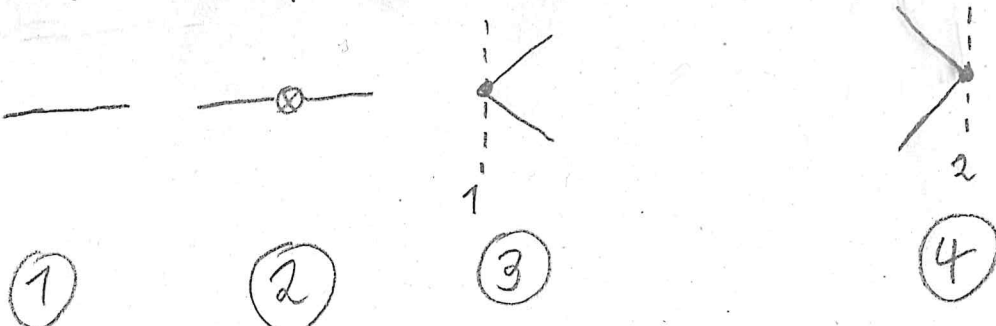
$$A = \Delta + \epsilon B \sim A^{-1} = \Delta^{-1} - \epsilon \Delta^{-1} B \Delta^{-1} + \mathcal{O}(\epsilon^2)$$

$$\text{check: } AA^{-1} = 1 + \epsilon (B \Delta^{-1} - \underbrace{\Delta \Delta^{-1}}_1 B \Delta^{-1}) + \mathcal{O}(\epsilon^2)$$

$$= 1 + \mathcal{O}(\epsilon^2)$$

$$\mathcal{D}_\phi^{-1} = \left( -\frac{\Delta}{4\pi G} + \frac{\partial_t^2}{4\pi G c^2} + \frac{m_1}{c^2} \delta_1 + \frac{m_2}{c^2} \delta_2 \right)^{-1}$$

$$= -4\pi G \Delta^{-1} - \frac{4\pi G}{c^2} \Delta^{-1} \partial_t^2 \Delta^{-1} - \frac{(4\pi G)^2}{c^2} m_1 \Delta^{-1} \delta_1 \Delta^{-1} - \frac{(4\pi G)^2}{c^2} m_2 \Delta^{-1} \delta_2 \Delta^{-1} + \mathcal{O}(c^{-4})$$





$$\gamma_\phi = -m_1 \left(1 + \frac{3V_1^2}{2c^2}\right) \delta_1 - m_2 \left(1 + \frac{3V_2^2}{2c^2}\right) \delta_2$$

contributions from ①:

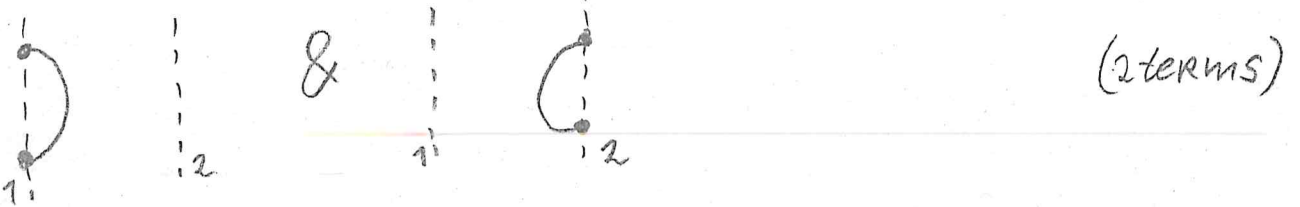
$$= \frac{1}{2} \int d^3X \left[ \gamma_{\phi 1} (-4\pi G \Delta^{-1}) \gamma_{\phi 2} + \gamma_{\phi 2} (-4\pi G \Delta^{-1}) \gamma_{\phi 1} \right] \quad (2 \text{ terms})$$

2 id. terms

$$= \frac{2}{2} \int d^3X (-4\pi G) \gamma_{\phi 1} \Delta^{-1} \gamma_{\phi 2}$$

$$= \dots = \frac{G m_1 m_2}{r} \left(1 + \frac{3V_1^2}{2c^2} + \frac{3V_2^2}{2c^2}\right) + O(c^{-4})$$

divergent self-int.  $\rightarrow$  drop



FROM ②: 2 id. terms

$$= \frac{2}{2} \int d^3X \gamma_{\phi 1} \frac{(-4\pi G)}{c^2} \Delta^{-1} \partial_t^2 \Delta^{-1} \gamma_{\phi 2}$$

$$= \int d^3X \frac{4\pi G}{c^2} (\partial_t \gamma_{\phi 1}) (\partial_t \Delta^{-2} \gamma_{\phi 2}) + \text{(total time der.)}$$

drop

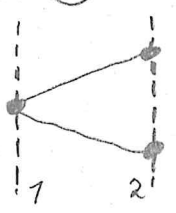
$$= \dots = \frac{G}{2rc^2} (V_1^i V_2^i - V_1^i n^i V_2^j n^j) + O(c^{-4})$$

$\Delta^{-2} \delta_2 = -\frac{1}{8\pi} |\vec{X} - \vec{X}_2(t)|$

$n^i = \frac{X_1^i - X_2^i}{r}$



from ③:



part. int.

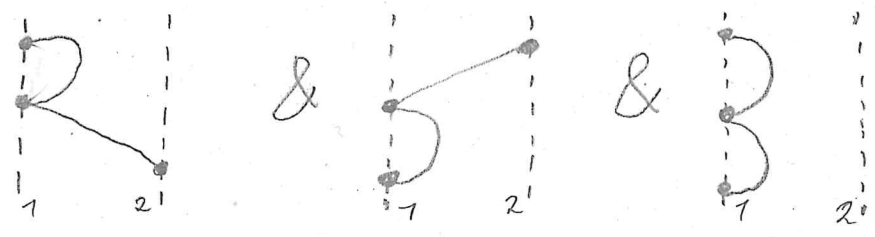
$$= \frac{1}{2} \int d^3x \int \phi_2 \left( -\frac{(4\pi G)^2}{c^2} m_1 \Delta^{-1} \delta_1 \Delta^{-1} \right) \int \phi_2$$

"symmetry factor"  $\approx -m_2 \delta_2$   $\approx -m_2 \delta_2$

$$= -\frac{G^2 m_1^2 m_2}{2c^2 r^2}$$

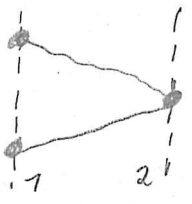
(1 term)

drop:



(3 terms)

from ④:



$$= -\frac{G^2 m_1^2 m_2}{2c^2 r^2}$$

analogous to ③  
& drop 3 terms

(1+3 terms)

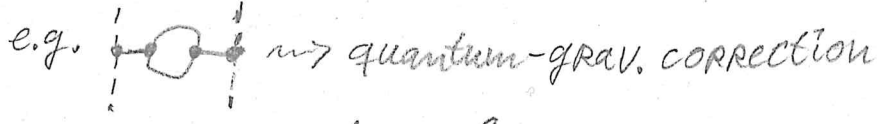
(total: 16 terms)

add:  $L_{kin} + L_A + L_\phi = \text{sum of diagrams}$

$\hookrightarrow$  eq. (2) on exercise sheet 3

connection to quantum field theory (QFT) and effective field theory

QFT: allow for all possible diagrams



derivation: e.g. via path integral

$\hookrightarrow$  "integrate out"  $g_{\mu\nu}$  from  $S[g_{\mu\nu}, X^a]$   
(classical part: Fokker action)

EFT:

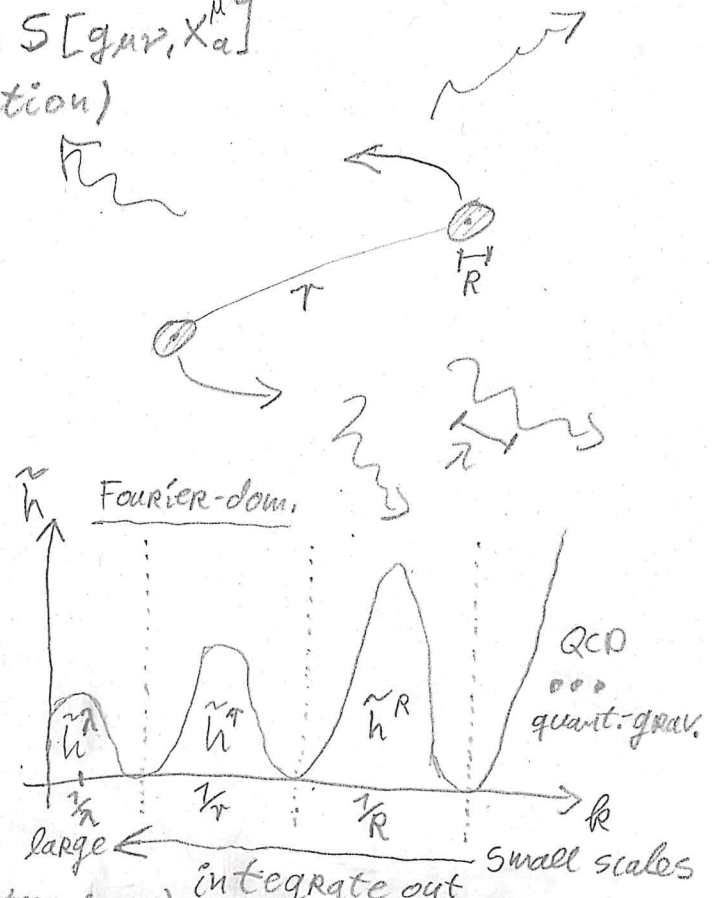
split:  $h_{\mu\nu} = g_{\mu\nu} - \eta_{\mu\nu}$   
 $= h_{\mu\nu}^{\lambda} + h_{\mu\nu}^{\tau} + h_{\mu\nu}^R + \dots$

according to hierarchy of scales

$\lambda > \tau > R > \dots$

$\hookrightarrow$  decoupling of Fourier modes:

$\tilde{h}^{\lambda} \cdot \tilde{h}^{\tau} = 0$   
 $\tilde{h}^{\lambda} \cdot \tilde{h}^R = 0$   
 $\tilde{h}^{\tau} \cdot \tilde{h}^R = 0$   
 $\dots$



EFT approach:

successively integrate out (path integral over) small scales  $\rightarrow$  coarse-graining, averaging

order:  $h_{\mu\nu}^R \rightarrow h_{\mu\nu}^{\tau} \rightarrow h_{\mu\nu}^{\lambda}$

integrate out  $h_{\mu\nu}^R$ :  $\rightarrow$  by Fokker action

$\hookrightarrow$  explicit calculation impossible due to strong gravity

$\hookrightarrow$  make ansatz based on

Symmetries & power counting  
(general cov.)  $(\frac{R}{\tau},$  multipole expansion)

$$S^{PP} = - \int d\tau \left( m c^2 + \frac{1}{2} \Omega^{\mu\nu} S_{\mu\nu} + \frac{1}{2} E_{\mu\nu} Q_E^{\mu\nu} + \frac{1}{2} B_{\mu\nu} Q_B^{\mu\nu} + \mathcal{O}\left(\frac{R^4}{\tau^4}\right) \right)$$

$\downarrow$  const. mass     $\downarrow$  angular vel.     $\downarrow$  spin     $\downarrow$  electric curvature component     $\downarrow$  gravito-electric quadrupole     $\downarrow$  magnetic curvature comp.     $\downarrow$  gravito-magnetic quadrupole

+ need model for multipoles

e.g. adiabatic tides:  $Q_E^{\mu\nu} = \text{const} \cdot E^{\mu\nu}$   
 $\hookrightarrow$  match to numerics